

Lecture 2 - Post-Newtonian Limit in GRI. Newtonian Gravity

$$\nabla^2 U = -4\pi\rho \quad \text{Poisson}$$

$$\rho \frac{dv}{dt} = \rho \nabla U - \nabla p \quad \text{Euler eq.}$$

$$\frac{d}{dt} = \frac{\partial}{\partial t} + \underline{v} \cdot \underline{\nabla}$$

$$\frac{\partial \rho}{\partial t} + \underline{\nabla} \cdot (\rho \underline{v}) = 0$$

$$\underline{v}^i = v^i(x, t)$$

$$\rho = \rho(p, T, Z, \dots)$$

$$U(x, t) = \int \frac{\rho(x', t)}{|x-x'|} d^3x'$$

Properties

$$\frac{1}{|x-x'|} = \sum_{lm} \frac{4\pi}{2l+1} \frac{r^l}{r'^l} Y_{lm}^*(\hat{n}_x) Y_{lm}(\hat{n}_{x'}) \quad \hat{n} = \frac{x}{r}$$

or, if  $|x'| < |x|$ ,

$$\frac{1}{|x-x'|} = \sum_{q=0}^{\infty} \frac{(-)^q}{q!} x'^q \nabla_{\varphi} \left( \frac{1}{r} \right)$$

$$x^{\varphi} \equiv x^i x^j \dots x^l$$

$$\equiv x^{ij\dots}$$

$$\nabla_{\varphi} \equiv \nabla_i \nabla_j \dots \nabla_l$$

$$= \nabla_{ij\dots}$$

can show

$$\nabla_{\varphi} \left( \frac{1}{r} \right) = (-)^q (2q-1)!! \frac{\hat{n}^{\langle \varphi \rangle}}{r^{q+1}}$$

$$n^{\langle \varphi \rangle} = \text{STF}$$

$$n^{\langle ij \rangle} = n^{ij} - \frac{1}{3} \delta^{ij}$$

$$n^{\langle ijkl \rangle} = n^{ijkl} - \frac{1}{5} (\delta^{ij} n^{kl} + \delta^{il} n^{jk} + \delta^{ik} n^{jl})$$

link between  $\hat{n}^{(q)}$  and  $Y_{em}$

$$N^q \hat{n}^{(q)} = \frac{q!}{(2q-1)!!} P_q(\cos\theta)$$



recall

$$P_\ell(\hat{N} \cdot \hat{n}) = \frac{4\pi}{2\ell+1} \sum_{m=-\ell}^{\ell} Y_{\ell m}^*(N) Y_{\ell m}(n)$$

Prove equivalence of 2 forms of  $1/|x-x'|$

$$\int \hat{n}^q d\Omega = \begin{cases} 0 & q \text{ odd} \\ \frac{4\pi}{(2q+1)!!} \left( \underbrace{\delta^{ij} \delta^{kl} \dots}_{q/2} + (\text{sym}) \right) \end{cases}$$

Equation of continuity

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho v) = 0$$

[assume no flow out of region]

$$\frac{\partial}{\partial t} \int \rho(x',t) f(x,x',t) d^3x' = \int \rho(x',t) \left[ \frac{\partial f}{\partial t} + v' \cdot \nabla' f(x,x',t) \right] d^3x'$$

$$\frac{d}{dt} \int \rho(x',t) f(x,x',t) d^3x' = \int \rho(x',t) \left[ \frac{\partial f}{\partial t} + v \cdot \nabla f + v' \cdot \nabla' f \right] d^3x'$$

eg show PPT

$$\frac{d}{dt} \int \rho(x',t) x'^i d^3x' = \int \rho(x',t) v'^i d^3x'$$

$$\dot{U} = \frac{\partial U}{\partial t} = \int \rho(x',t) \frac{v' \cdot (x-x')}{|x-x'|^3} d^3x'$$

$$\frac{dU}{dt} = - \int \rho(x',t) \frac{(v-v') \cdot (x-x')}{|x-x'|^3} d^3x'$$

## Euler Equation

$$\rho \frac{dv}{dt} = \rho \nabla U - \nabla p \quad v = v(x, t)$$

Integrate over all space

$$\int \rho \frac{dv^i}{dt} = \frac{d}{dt} \int \rho v^i d^3x = \frac{d^2}{dt^2} \int \rho v^i d^3x$$

$$\int \rho \nabla^i U d^3x = - \int \rho(x, t) \int \rho(x', t) \frac{(x-x')^i}{|x-x'|^3} d^3x d^3x' \equiv 0$$

$$\int \nabla^i p d^3x = \int p n^i R^2 d\Omega = 0$$

Define  $M = \int \rho d^3x$      $X = \frac{1}{M} \int \rho x$  ,     $V = \frac{1}{M} \int \rho v$

$$\therefore \boxed{\frac{d}{dt} (MV^i) = \frac{d^2}{dt^2} (MX^i) = 0 \quad \text{cons. of momentum}}$$

$$\int \rho v \cdot \frac{dv}{dt} = \int \rho v \cdot \nabla U - \int v \cdot \nabla p$$

$$\textcircled{1} = \frac{1}{2} \frac{d}{dt} \int \rho v^2 d^3x \equiv \frac{d}{dt} \mathcal{T}$$

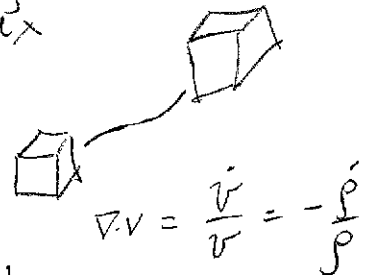
$$\textcircled{2} = - \int \rho \rho' \frac{v \cdot (x-x')}{|x-x'|^3} d^3x d^3x' = - \frac{1}{2} \int \rho \rho' \frac{(v-v') \cdot (x-x')}{|x-x'|^3}$$

$$= + \frac{1}{2} \frac{d}{dt} \int \frac{\rho \rho'}{|x-x'|} d^3x d^3x' \equiv - \frac{d}{dt} \Omega$$

$$\textcircled{3} = - \int v \cdot \nabla p = - \int \nabla \cdot (vp) d^3x + \int \rho \nabla \cdot v d^3x$$

$$= \int \frac{\rho dv/dt}{v} d^3x$$

$$= - \left[ \frac{1}{\rho} \frac{d}{dt} (\rho \Pi / v) \right] = - \left( \rho \frac{d\Pi}{dt} \right) = - \frac{d}{dt} (\rho \Pi) d^3x$$



Thus

$$\boxed{\mathcal{L} + \mathcal{L} + \mathcal{E} = \text{const}}$$

### Newtonian Stress Energy Tensor

$$T^{00} = \rho$$

$$T^{0j} = \rho v^j$$

$$T^{ij} = \rho v^i v^j + p \delta^{ij} + \frac{1}{4\pi} \left( U^{ij} U^{ij} - \frac{1}{2} \delta^{ij} U^{jk} U^{jk} \right)$$

Then

$$T^{0\nu}_{;\nu} = 0 \Rightarrow \text{continuity eq'n}$$

$$T^{j\nu}_{;\nu} = 0 \Rightarrow \text{Euler's eq'n}$$

$$\boxed{T^{\mu\nu}_{;\nu} = 0}$$

$\boxed{\text{PPT}}$

Equations of Motion

For a system of bodies, size  $R$ , separation  $r$ ,

$$U_A = \sum_{q=0}^{\infty} \frac{(-1)^q}{q!} I_A^{(q)} \nabla \phi \left( \frac{L}{r} \right) \quad r = |x - x_A|$$

$$I_A^{(q)} = \int_A \rho' \bar{x}'^q d^3x' \approx m R^q \quad \bar{x}' = |x' - x_A|$$

$$I_A^{(0)} = m \quad I_A^{(1)} = 0$$

So for  $R \ll r$ , only the  $q=0$  term counts

$$U_A = \frac{M_A}{|x - x_A|}, \quad \int_A \nabla \phi d^3x = 0$$

$$M_A a_A^i = - \sum_{B \neq A} \frac{M_A M_B (x - x_B)^i}{|x - x_B|^3}$$

Two body motion

$$a_1^i = - \frac{M_2 x^i}{r^3} \quad x = |x_1 - x_2|$$

$$a_2^i = \frac{M_1 x^i}{r^3} \quad v = v_1 - v_2$$

$$a = a_1 - a_2$$

Relative equation of motion

$$\boxed{\ddot{a} = - \frac{m x}{r^3}} \quad m = m_1 + m_2$$

$$E = \frac{1}{2} v^2 + U \quad U = \frac{m}{r}$$

## 2. Post-Newtonian Theory

gravity weak  $\sim U \sim m/r \sim 10^{-5} - 10^{-8}$  in SS

slow motion  $V^2 \sim U$  -

pressure low  $\frac{p}{\rho} \sim V^2 \sim U$  -

counting orders for dimensionless variables or functions

$$\epsilon \sim V^2 \sim U \sim p/\rho \quad \epsilon^{1/2} \sim v \sim \frac{d/dt}{d/dx}$$

Geodesics

$$\delta S = 0$$

$$S = \int \sqrt{-g_{00} + 2g_{0i}v^i + g_{ij}v^i v^j} dt$$

NGT?  $\quad \quad \quad L$

- recall  $L = -\frac{1}{2}V^2 - U$

$$L_R = \left( 1 - \delta g_{00} - 2g_{0i}v^i - V^2 - \delta g_{ij}v^i v^j \right)^{1/2}$$

$$\approx 1 - \frac{1}{2}\delta g_{00} - \frac{1}{2}V^2 + \dots$$

If  $\delta g_{00} = +2U$ , get NGT

Note  $U \sim \epsilon \sim V^2$

What about next corrections

$$L_R = \left( 1 - 2U - \delta g_{00}^{(2)} - 2\delta g_{0i}^{(1.5)} v^i - V^2 - \delta g_{ij}^{(1)} v^i v^j \right)^{1/2}$$

$\epsilon \quad \quad \quad \epsilon^2 \quad \quad \quad \epsilon^{3/2} \quad \epsilon^{1/2} \quad \quad \quad \epsilon \quad \quad \quad \epsilon \quad \quad \quad \epsilon$

Therefore need

$$g_{00} \text{ to } O(\epsilon^2)$$

$$g_{0i} \text{ to } O(\epsilon^{3/2})$$

$$g_{ij} \text{ to } O(\epsilon)$$

What about light?

### 3. PN Limit of GR

$$G_{\mu\nu} = 8\pi T_{\mu\nu} \quad \left[ \begin{array}{l} \text{Bianchi identities} \\ G^{\alpha\beta}_{;\beta} = 0 \Rightarrow T^{\alpha\beta}_{;\beta} = 0 \end{array} \right]$$

weak fields everywhere so that  $g_{\mu\nu} \sim \eta_{\mu\nu}$

Define

$$g^{\mu\nu} = \sqrt{-g} g^{\mu\nu} \quad g = \det(g_{\mu\nu})$$

$$H^{\mu\alpha\nu\beta} = g^{\mu\nu} g^{\alpha\beta} - g^{\alpha\nu} g^{\mu\beta} \quad - \text{symmetries of Riem}$$

can prove

$$\begin{aligned} H^{\mu\alpha\nu\beta}_{;\mu\nu} &= (-g)(2G^{\alpha\beta} + 16\pi t^{\alpha\beta}_{LL}) \\ &= 16\pi(-g)T^{\alpha\beta} + 16\pi(-g)t^{\alpha\beta}_{LL} \end{aligned}$$

$$t^{\alpha\beta}_{LL} \sim (\partial_\gamma g^{\alpha\beta})^2$$

Now define  $g^{\mu\nu} = \eta^{\mu\nu} - h^{\mu\nu}$

Coordinate freedom: choose

$$h^{\mu\nu}_{;\nu} = 0 \quad \text{Lorentz, deDonder, Harmonic coords}$$

Then  $\left[ \text{Ex: Prove } \square X^\alpha = 0 \quad \square = \frac{1}{\sqrt{g}} \partial_\alpha (\sqrt{g} g^{\alpha\beta} \partial_\beta) \right]$

$$H^{\mu\alpha\nu\beta}_{;\mu\nu} = -\square_\gamma h^{\alpha\beta} + h^{\alpha\beta}_{;\mu\nu} h^{\mu\nu} - h^{\alpha\nu}_{;\mu} h^{\mu\beta}_{;\nu}$$

Relaxed E. Eq.

$$\begin{aligned} \square_\gamma h^{\alpha\beta} &= -16\pi(-g)T^{\alpha\beta} - 16\pi(-g)t^{\alpha\beta}_{LL} + h^{\alpha\beta}_{;\mu\nu} h^{\mu\nu} - h^{\alpha\nu}_{;\mu} h^{\mu\beta}_{;\nu} \\ &= \dots - T^{\alpha\beta} \end{aligned}$$



Note  $T^{\alpha\beta}_{;\beta} = 0 \iff T^{\alpha\beta}_{;\beta} = 0$

Define

$h^{00} = N$   $h^{0i} = K^i$   $h^{ij} = B^{ij}$   $h^{ic} = B$   
 $\epsilon$   $\epsilon^{3/2}$   $\epsilon^2$   $\epsilon^2$

Invert

$g_{00} = -\left(1 - \frac{1}{2}N + \frac{3}{8}N^2\right) + \frac{1}{2}B + O(\epsilon^3)$

$g_{0i} = -K^i + O(\epsilon^{5/2})$

$g_{ij} = \delta^{ij}\left(1 + \frac{1}{2}N\right) + O(\epsilon^2)$

N to  $O(\epsilon)$

Recall:  $T^{\alpha\beta} = (\rho + p)u^\alpha u^\beta + pg^{\alpha\beta}$

$\square_{\eta} N_0 = -16\pi\rho$   
"  $\nabla^2 N$

$N_0 = 4U$

$u^\alpha = \frac{dx^\alpha}{dt}$   
 $= \frac{dt}{dt}(1, \underline{v})$

B to  $O(\epsilon^2)$

$\nabla^2 B = -16\pi T^{00}$   
 $= -16\pi(\rho v^2 + 3\rho) + \frac{1}{8}\nabla N_0^2$   
 $= -16\pi(\rho v^2 + 3\rho) + 2(\nabla U)^2$   
 $= -16\pi(\rho v^2 + 3\rho) + \nabla^2(U^2) + 8\pi\rho U$

$B = 4\Phi_1 + 12\Phi_4 + U^2 - 2\Phi_2$

$\rho\rho T$

$K^i$  to  $O(\epsilon^{3/2})$

$$\nabla^2 K^i = -16\pi \rho V^i$$

$$K^i = 4V^i$$

$N$  to  $O(\epsilon^2)$

$$\square_\gamma N = -16\pi(-g)T^{00} + \frac{7}{8}|\nabla N_0|^2$$

$$(-g) \approx (1 - \frac{1}{2}N_0)(1 + \frac{1}{2}N_0)^3 = 1 + 4U$$

$$T^{00} = [\rho(1+\pi) + p]u^0{}^2 - p$$
$$= \rho(1 + 2u + v^2 + \pi)$$

$$|\nabla N_0|^2 = 16|\nabla u|^2 = 8 \cdot \nabla^2(U^2) + 64\pi\rho U$$

$$\square_\gamma N = \nabla^2 N - \ddot{N}_0$$
$$= \nabla^2 N - 4\ddot{U}$$
$$= \nabla^2(N - 2\ddot{X})$$

$$U = \frac{1}{2}\nabla^2 X$$

↳ superpotential

$$\left(\nabla^2|x-x'| = \frac{2}{|x-x'|}\right)$$

$$\nabla^2(N - 2\ddot{X}) = -16\pi\rho - 96\pi\rho U - 16\pi\rho v^2 - 16\pi\rho\pi$$
$$+ 7\nabla^2 U^2 + 56\pi\rho U$$

$$N = 4U + 10\Phi_2 + 4\Phi_1 + 4\Phi_3 + 7U^2 + 2\ddot{X}$$

### 4. Gauge transf<sup>n</sup>

- general covariance - any coords can be used
- 2 coord transf<sup>n</sup> - PN gauge
  - Lorentz transf

$$g_{\bar{\alpha}\bar{\beta}}(\bar{x}, \bar{t}) = \frac{\partial x^\mu}{\partial \bar{x}^\alpha} \frac{\partial x^\nu}{\partial \bar{x}^\beta} g_{\mu\nu}(x, t)$$

$$x^{\bar{\alpha}}(x^\mu)$$

#### A. PN gauge

$$x^{\bar{\mu}} = x^\mu + \xi^\mu(x^\nu) \quad \frac{|\xi|}{|x|} \ll 1 \quad [\text{first order in } \xi]$$

$$\frac{\partial x^\mu}{\partial \bar{x}^\alpha} = \delta_\alpha^\mu - \xi_{,\alpha}^\mu$$

$$g_{\bar{\alpha}\bar{\beta}}(\bar{x}) = g_{\alpha\beta}(x) - \xi_{\alpha;\beta} - \xi_{\beta;\alpha}$$

$\xi_{(\alpha;\beta)} \sim$  PN potential,  $\xi_\alpha$  units of distance

$$\xi_{(\alpha;\beta)} \rightarrow 0 \text{ at } \infty \quad \frac{|\xi^{\alpha\beta}|}{|x^\alpha|} \rightarrow 0 \text{ at } \infty$$

$$\xi_{0;0} \sim \epsilon^2 \quad \xi_{0;j} \sim \epsilon^{3/2}, \quad \xi_{i;j} \sim \epsilon$$

$$\Rightarrow \xi_0 = \lambda_1 X_{10}, \quad \xi_j = \lambda_2 X_{1j}$$

$$g_{\bar{0}\bar{0}} = g_{00} + 2\lambda_2 X_{1;0} + O(\epsilon^2)$$

$$g_{\bar{0}\bar{j}} = g_{0j} - (\lambda_1 + \lambda_2) X_{1;0j} + O(\epsilon^{5/2})$$

Also, in  $g_{00}$  we have

$$U(\bar{x}, \bar{t}) = \int \frac{\rho(x', \bar{t})}{|\bar{x} - x'|} d^3x'$$

- need to convert  $x'$  to  $\bar{x}'$

$$\rho(x', \bar{t}) = \rho(\bar{x}, \bar{t}) \text{ - invariant - measured in LFF}$$

$$\sqrt{-g} u^0 d^3x = \text{invariant} : d^3x' \rightarrow d^3\bar{x}' (1 + 2\lambda_2 u)$$

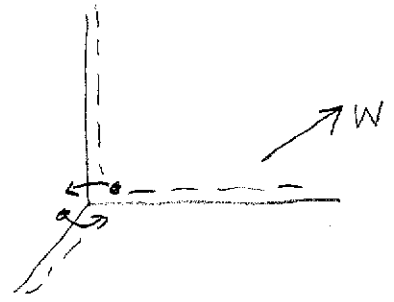
$$\frac{1}{|\bar{x} - x'|} = \frac{1}{|\bar{x} - \bar{x}'|} + \nabla' \frac{1}{|\bar{x} - \bar{x}'|} \cdot \frac{x'}{\bar{x} - \bar{x}'}$$

Final result PPT  
PPT

B. Lorentz transf<sup>n</sup>

$$X^\alpha = \Lambda^\alpha_\beta \xi^\beta$$

$$\Lambda^\alpha_\beta = \begin{pmatrix} \gamma & -\gamma w \\ -\gamma w & \mathbb{I} + (\gamma-1)\hat{w}^i\hat{w}^j \end{pmatrix}$$



for small  $w$  ( $w \tau \sim O(\epsilon)$ )

$$X = \xi + (1 + \frac{1}{2}w^2)\underline{w}\tau + \frac{1}{2}(\xi \cdot \underline{w})\underline{w} + O(\epsilon^2)\xi$$

$$t = \tau(1 + \frac{1}{2}w^2 + \frac{3}{8}w^4) + (1 + \frac{1}{2}w)\underline{w} \cdot \xi + O(\epsilon^3)\tau$$

$$g_{\mu\nu}(\xi, \tau) = \frac{\partial X^\alpha}{\partial \xi^\mu} \frac{\partial X^\rho}{\partial \xi^\nu} g_{\alpha\rho}(x, t)$$

$$\bar{\rho} = \rho \text{ etc.} \quad v = v + w + O(\epsilon^{3/2})$$

$$d^3x' \rightarrow d^3\xi' (1 - v' \cdot w - \frac{1}{2}w^2) \quad - \text{change in LF contraction}$$

$$\frac{1}{|x-x'|} = \frac{1}{|\xi-\xi'|} \left( 1 + \frac{1}{2}(w \cdot \hat{n})^2 + (w \cdot \hat{n})(v' \cdot \hat{n}) + \dots \right)$$

$\uparrow$  same  $\uparrow$  same  
 $t$   $\tau$

Possible additional gauge transf<sup>n</sup>

$$\bar{t} = \tau + \lambda_3 \underline{w} \cdot \underline{\nabla} X$$