# **Domain Walls as Probes of Gravity**

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High Energy, Cosmology and Strings Institut Henri Poincaré, December 2006 What: Gravitational effect of Domain Walls in IR-modified gravity

Why: Gravitating DW solutions in DGP can be found **exactly**. Useful information about intrinsically non-linear theory.

Main Result 1) Screening of 4D and 5D tension.

Analog to screening in Schwarzchild-like solutions.

Main Result 2) Domain Walls are short-distance probes of large-distance modified gravity.

### Dvali Gabadadze Porrati (DGP)

$$S = \int d^5x \sqrt{-g} \, \frac{M_*^3}{2} \, R_5 + \int d^4x \sqrt{-h} \, \frac{m_P^2}{2} \, R_4$$

$$r_c = \frac{m_P^2}{2M_*^3}$$

bulk

$$G_{AB}^{(5)} = 0 \qquad \left(x_5 \neq 0\right)$$

$$M_*^3 \left( K_{\mu\nu} - K h_{\mu\nu} \right) = T_{\mu\nu} - m_P^2 G_{\mu\nu}^{(4)}$$

brane

### vDVZ discontinuity

IR modified

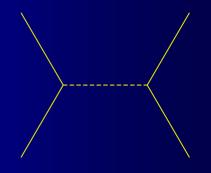
(more d.o.f.'s)

$$\mathcal{A} \propto G_N \frac{T_{\mu\nu}T'^{\mu\nu} - \frac{1}{3}TT'}{\Box - m^2(\Box)}$$

$$m^2(\Box) = \sqrt{\Box}/r_c$$

GR

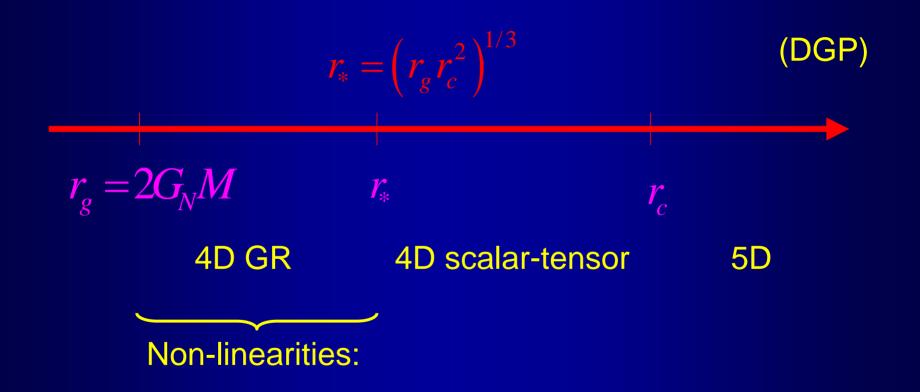
$$\mathcal{A} \propto G_N \frac{T_{\mu\nu}T'^{\mu\nu} - \frac{1}{2}TT'}{\Box}$$

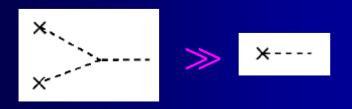


way out:

higher orders in Perturbation Theory become important at short distances

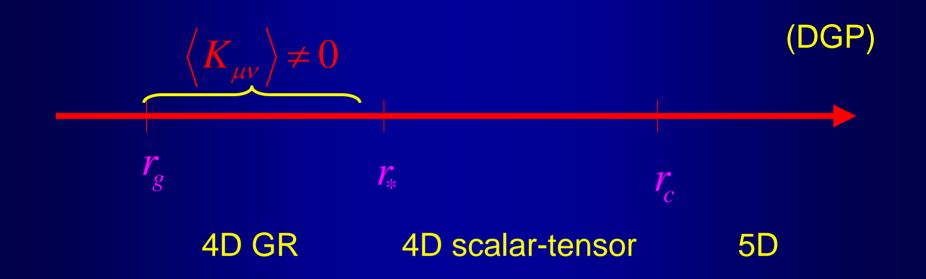
#### Non-perturbative continuity





Weak field expansion breaks down. Large  $r_c$  expansion is OK.

### Non-perturbative continuity

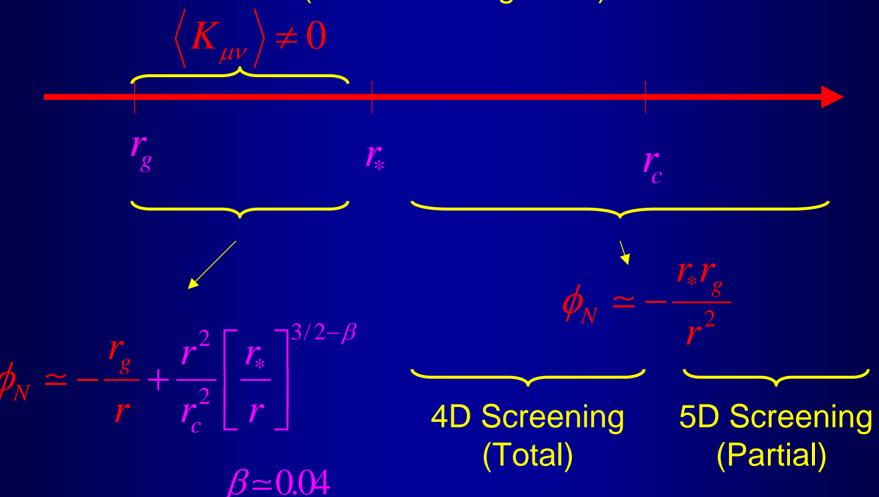


$$\phi_N \simeq -\frac{r_g}{r} + \frac{r^2}{r_c^2} \left[ \frac{r_*}{r} \right]^{3/2} \qquad -\alpha \frac{r_g}{r} \qquad -\beta \frac{2G_5 M}{r^2}$$

Schwarzchild-AdS like

#### 'Schwarzchild-DGP'

(Gabadadze-Iglesias)



#### 'Schwarzchild-DGP'

(Gabadadze-Iglesias)

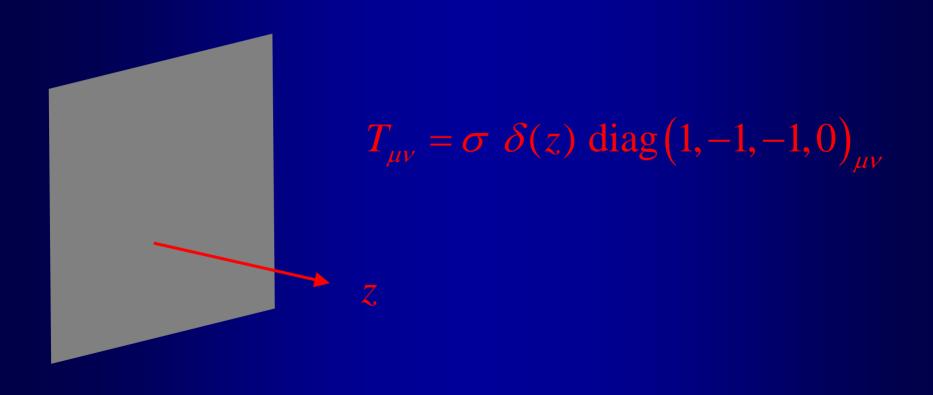
$$\phi_N \simeq -\frac{2G_4M^{(4D)}}{r} - \frac{2G_5M^{(5D)}}{r^2}$$

#### Screening:

branch	$M^{(4D)}/M$	$M^{(5D)}/M$
СВ	0	$0.56 \; \frac{r_*}{r_c}$
SAB	0	$-0.45 \frac{r_*}{r_c}$

#### OUTLINE

- Domain Walls in GR
- Strings in GR (sub-critical, super-critical)
- Domain Walls on a DGP brane
  - Conventional Branch
  - Self-Accelerated Branch
- Screening of the tension (both 4D and 5D)
- Extension to other IR-modified gravity



$$ds^{2} = (1 - H |z|)^{2} (-dt^{2} + e^{2Ht} [dx^{2} + dy^{2}]) + dz^{2}$$

$$H = \frac{\sigma}{4m_P^2}$$
 horizon

DW inflates.

Hubble rate= H



Vilenkin-Ipser-Sikivie

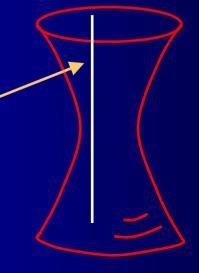
$$ds^{2} = (1 - H | z |)^{2} (-dt^{2} + e^{2Ht} [dx^{2} + dy^{2}]) + dz^{2}$$

$$H = \frac{\sigma}{4m_P^2}$$
 horizon

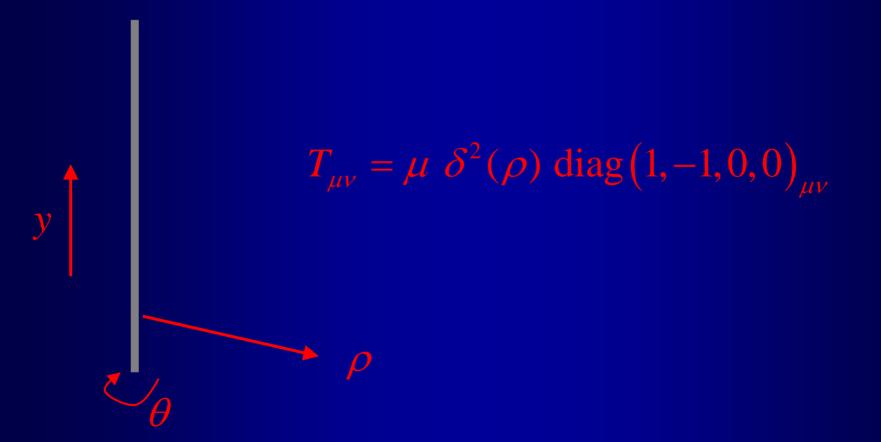
Repulsive force



geodesic



# Strings in GR (co-dimension=2)



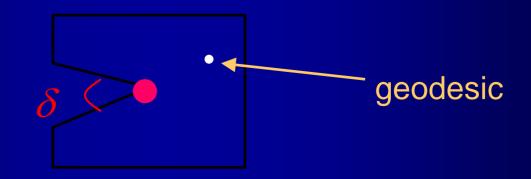
# Strings in GR (co-dimension=2)

$$ds^{2} = \left(-dt^{2} + dy^{2}\right) + d\rho^{2} + \left(1 - \frac{\mu}{m_{P}^{2}}\right)^{2} \rho^{2} d\theta^{2}$$

$$\delta = 2\pi \frac{\mu}{m_p^2}$$
 D

Deficit angle

No force (on static particles)



# Strings in GR (co-dimension=2)

$$\mu = 2\pi m_P^2$$
 'critical string'

$$\delta = 2\pi$$

$$ds^2 = \left(-dt^2 + dx^2\right) + d\rho^2 + D^2d\theta^2$$

D =thickness

Transverse space compactified.



(flat worldsheet)

### Strings in GR (co-dimension=2)

$$\mu > 2\pi m_P^2$$

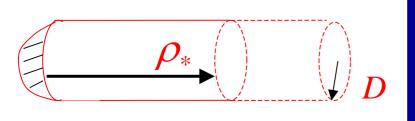
'super-critical string'

$$\delta = 2\pi$$



(Transverse space compactified)

Topological Inflation 
$$H_* \simeq \frac{\mu - 2\pi m_P^2}{4D m_P^2}$$



Horizon at 
$$\rho_* = \frac{1}{H_*}$$

$$\frac{R_0}{r} \sim \frac{\sigma}{M_*^3}$$

Critical tension  $\sigma_c \sim M_*^3$ 

$$\sigma_c \sim M_*^3$$

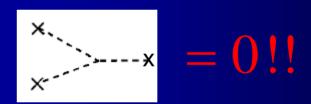
Deficit angle

$$\delta \sim 2\pi$$

# Domain Walls in DGP: Perturbative arguments

What is r for a DW??

**~** 



$$\left(\partial\chi\right)^2$$

$$(\partial \chi)^2 r_c^2 \partial^2 \chi$$

$$(z \gg r_*)$$

$$\Box \chi = \frac{T}{m_P} = \frac{\sigma}{m_P} \delta(z)$$

$$\chi \propto |z|$$

$$r_* = d$$
 thickness

Domain wall with maximal symmetry



Flat

 $\kappa = 0$ 

(2+1) De Sitter

 $\kappa = 1$ 

Birkhoff Theorem:

$$ds^2 = f(R)dZ^2 + \frac{dR^2}{f(R)} + R^2 ds_{\kappa}^2$$

$$f(R) = \kappa - C/R^2$$

Non-singular



C = 0

5D Minkowski bulk

Non-compact bulk

$$ds^2 = dZ^2 + dR^2 + R^{2\kappa} ds_{\kappa}^2$$

Cartesian  $\kappa = 0$ 

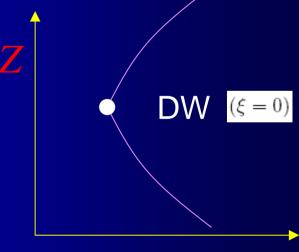
Brane parametrization (embedding):

$$(R(\xi), Z(\xi))$$

$$Z^{\prime 2} + R^{\prime 2} = 1$$

Induced metric on the brane:

$$ds_4^2 = d\xi^2 + R^{2\kappa}(\xi) \ ds_\kappa^2$$



Equations for the embedding: Israel junction conditions

$$M_*^3 \left( K_{\mu\nu} - K h_{\mu\nu} \right) = T_{\mu\nu} - m_P^2 G_{\mu\nu}^{(4)}$$

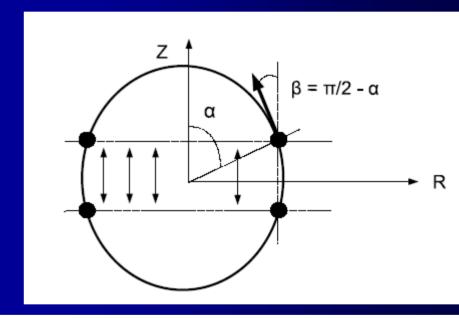
A) 
$$\epsilon \kappa \frac{\sqrt{1 - R'^2}}{r_c R} = -\kappa \frac{1 - R'^2}{R^2} + \frac{\tau}{3m_P^2}$$

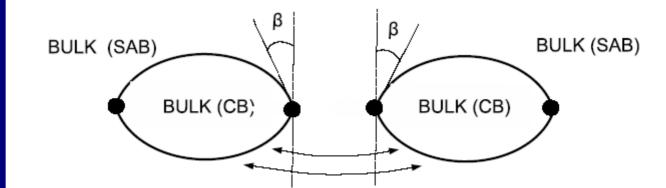
$$-\epsilon \frac{R''}{2r_c\sqrt{1-R'^2}} = \kappa \frac{R''}{R} + \frac{\sigma}{2m_P^2}\delta(\xi)$$

#### A) (Outside the DW)

$$R(\xi) = H^{-1} \sin \left[ H(\xi_0 - |\xi|) \right]$$

$$\epsilon \kappa \frac{H}{r_c} = -\kappa H^2 + \frac{\tau}{3m_P^2}$$





B) Junction condition on the DW

$$-\kappa \frac{\Delta R_0'}{R_0} + \frac{\delta}{4r_c} = \frac{\sigma}{2m_P^2}$$

$$2\kappa \,\epsilon \,Hr_c \tan\frac{\delta}{4} \,+\, \frac{\delta}{4} = \frac{\sigma}{4M_*^3}$$

#### B) Junction condition on the DW

$$-2\kappa \ m_P^2 \ \frac{\Delta R_0'}{R_0} + \delta \ M_*^3 = \sigma$$

4D viewpoint:

$$\sigma_{\rm eff}^{(4D)} \equiv \sigma - \delta \, M_*^3$$

Extrinsic curvature

5D viewpoint:

$$\sigma_{\rm eff}^{(5D)} \equiv \sigma + \kappa \ m_P^2 \frac{\Delta R_0'}{R_0}$$

**DGP** term

$$\sigma_{\rm eff}^{(4D)} + \sigma_{\rm eff}^{(5D)} = \sigma$$

**Conventional Branch** 

Subcritical Wall  $\sigma < 2\pi M_*^3$ 

$$\sigma < 2\pi M_*^3$$

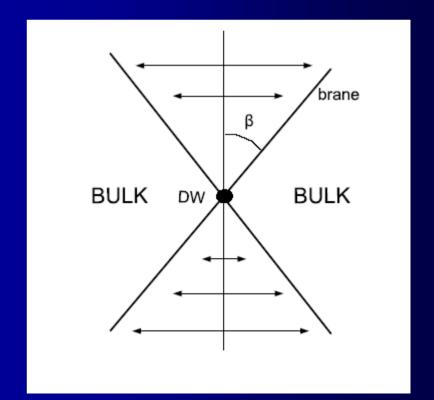
$$\kappa = 0$$

$$R(\xi) = |\xi| \sin \beta$$

$$Z(\xi) = \xi \cos \beta$$

$$4\beta = \delta = \frac{\sigma}{M_*^3}$$

'brane bending'



Conventional Branch

Subcritical Wall  $\sigma < 2\pi M_*^3$ 

$$\sigma < 2\pi M_*^3$$

$$\kappa = 0$$

$$\sigma_{\rm eff}^{(5D)} \equiv \sigma + \kappa \ m_P^2 \frac{\lambda R_0'}{R_0}$$

No 4D screening

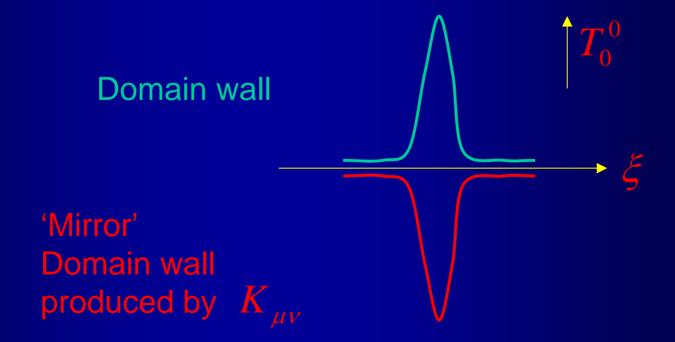
$$\sigma_{\text{eff}}^{(4D)} \equiv \sigma - \delta M_*^3 = 0!!$$

Absolute 4D screening

**Conventional Branch** 

Subcritical Wall  $\sigma < 2\pi M_*^3$ 

$$\sigma < 2\pi M_*^3$$



**Conventional Branch** 

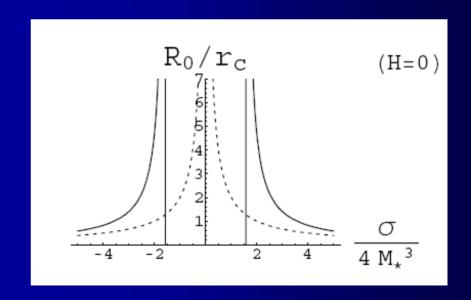
Super-critical Wall  $\sigma > 2\pi M_*^3$ 

$$\sigma > 2\pi M_*^3$$

$$\kappa = 1$$

$$\frac{1}{R_0} + \frac{\pi}{4r_c} = \frac{\sigma}{4m_P^2}$$

Partial 4D Screening

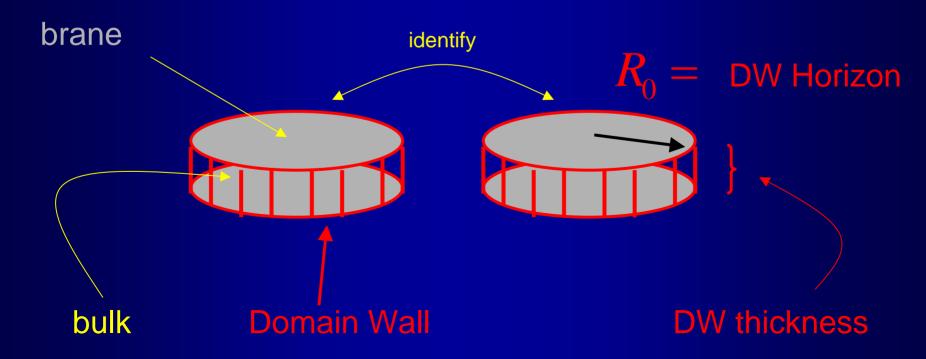


**Conventional Branch** 

geometry

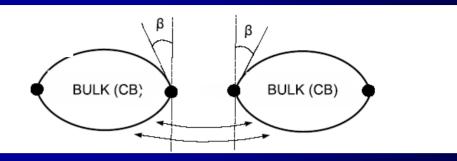
Super-critical Wall  $\sigma > 2\pi M_*^3$ 

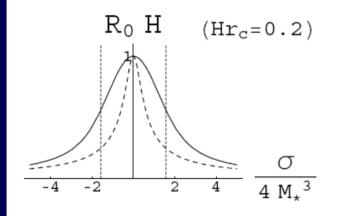
$$\sigma > 2\pi M_*^3$$

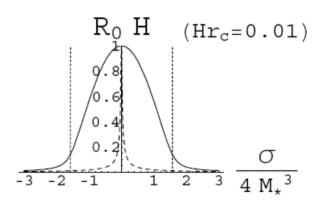


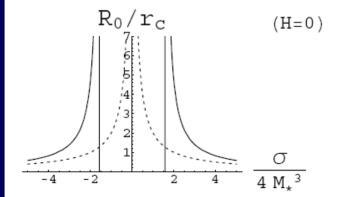
### Non-zero brane tension

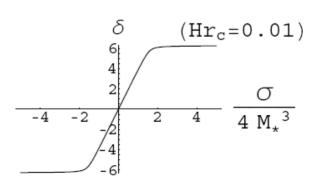
 $H \neq 0$ 











Self-Accelerated Branch

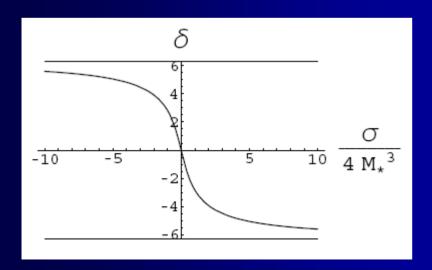
$$(\kappa = 1)$$

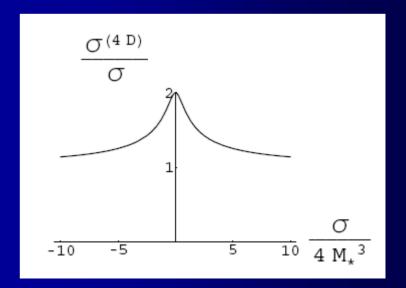
$$\delta < 0!!$$

5D 'over-screening'

$$\sigma_{eff}^{(4D)} > \sigma!!$$

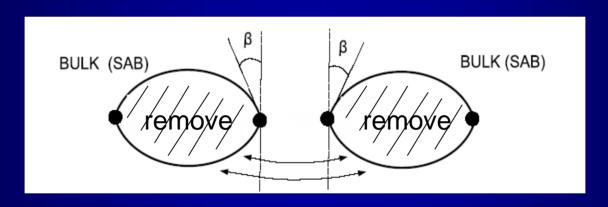
4D 'anti-screening'



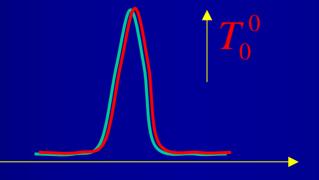


**Self-Accelerated Branch** 

geometry



Domain wall



'Replicant'
Domain wall
produced by



#### summary

branch	DW tension	$\sigma_{\mathrm{eff}}^{(4D)}/\sigma$	$\sigma_{ m eff}^{(5D)}/\sigma$
СВ	$\sigma < 2\pi M_*^3$	0	1
	$\sigma > 2\pi M_*^3$	$1 - \frac{2\pi M_*^3}{\sigma}$	??
SAB	$\sigma \ll 2\pi M_*^3$	$2\left(1-\left(\sigma/M_{*}^{3}\right)^{2}/48+\ldots\right)$	$-\left(1-\left(\sigma/M_{*}^{3}\right)^{2}/24+\ldots\right)$
	$\sigma\gg 2\pi M_*^3$	$1 + \frac{2\pi M_*^3}{\sigma} + \dots$	??

	=
Schwarzchild-DGP	_

		ı
branch	$M^{(4D)}/M$	$M^{(5D)}/M$
СВ	0	$0.56 \; \frac{r_*}{r_c}$
SAB	0	$-0.45 \frac{r_*}{r_c}$

### Super-critical walls and 5D screening

With No DGP 'induced gravity' term (supermassive codimension-2 brane)

$$\sigma > 2\pi M_*^3$$



$$T^{\nu}_{\mu} = \operatorname{diag}(-\rho, -\rho, -\rho, P)$$

$$2\pi M_*^3 = \int d\xi \, (\rho + 2P/3) \, \Box -\frac{2}{3} P_0 \, d = \sigma - 2\pi M_*^3$$

$$-\frac{2}{3}P_0 d = \sigma - 2\pi M_*^3$$

$$-P_0 = \frac{6M_*^3}{R_0}$$

$$\frac{1}{R_0} = \frac{1}{4dM_*^3} \left(\sigma - 2\pi M_*^3\right)$$

### Super-critical walls and 5D screening

With No DGP 'induced gravity' term (supermassive codimension-2 brane)

$$\sigma > 2\pi M_*^3$$



$$\frac{1}{R_0} = \frac{1}{4dM_*^3} \left(\sigma - 2\pi M_*^3\right)$$

With No DGP 'induced gravity' term,

$$\sigma > 2\pi M_*^3$$



$$\frac{1}{R_0} + \frac{\pi}{4r_c} = \frac{\sigma}{4m_P^2}$$

### Super-critical walls and 5D screening

With No DGP 'induced gravity' term (supermassive codimension-2 brane)

$$\sigma > 2\pi M_*^3$$



$$\frac{1}{R_0} = \frac{1}{4dM_*^3} \left( \sigma - 2\pi M_*^3 \right)$$

With No DGP 'induced gravity' term,

$$\sigma > 2\pi M_*^3$$



$$\frac{1}{R_0} = \frac{1}{4m_P^2} \left( \sigma - 2\pi M_*^3 \right)$$

$$\frac{1/R_0^{(DGP)}}{1/R_0^{(5D)}} = \frac{d}{2r_c}$$

The SAME suppression factor!!

$$r_* = d$$

#### DW

branch	DW tension	$\sigma_{ m eff}^{(4D)}/\sigma$	$\sigma_{ m eff}^{(5D)}/\sigma$
СВ	$\sigma < 2\pi M_*^3$	0	1
	$\sigma > 2\pi M_*^3$	$1 - \frac{2\pi M_*^3}{\sigma}$	$\left(\frac{r_*}{2r_c}\right)$
SAB	$\sigma \ll 2\pi M_*^3$	$2\left(1-\left(\sigma/M_{*}^{3}\right)^{2}/48+\ldots\right)$	$-\left(1-\left(\sigma/M_*^3\right)^2/24+\ldots\right)$
	$\sigma \gg 2\pi M_*^3$	$1 + \frac{2\pi M_*^3}{\sigma} + \dots$	$\left(-\frac{r_*}{2r_c}\right)$

Schwarzchild-DGF	

	branch	$M^{(4D)}/M$	$M^{(5D)}/M$
٠	СВ	0	$0.56 \frac{r_*}{r_c}$
	SAB	0	$(-0.45 \frac{r_*}{r_c})$

### Generic IR-modified gravity

#### Linearized Einstein's eqs:

$$\mathcal{E}^{\alpha\beta}_{\mu\nu}h_{\alpha\beta} - m^2(\Box)(h_{\mu\nu} - \eta_{\mu\nu}h) = -16\pi G_N T_{\mu\nu}$$

$$m^2(\Box) = r_c^{-2(1-\alpha)} \Box^{\alpha}$$

$$0 \le \alpha < 1$$

$$\alpha = 1/2$$
 DGP

$$\alpha = 0$$
 Fierz-Pauli Massive gravity

### Generic IR-modified gravity

#### Couples to conserved source as

$$h_{\mu\nu} = -16\pi G_N \frac{1}{\Box - m^2(\Box)} \left\{ T_{\mu\nu} - \frac{1}{3} \left( \eta_{\mu\nu} - \frac{1}{m^2(\Box)} \partial_{\mu} \partial_{\nu} \right) T \right\}$$

'pure gauge' form



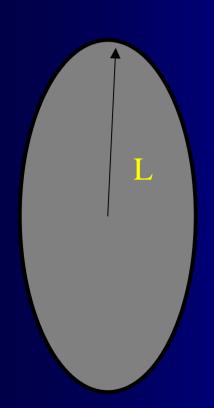
For DWs, 
$$T_{\mu\nu} - (1/3)T\eta_{\mu\nu} \propto \delta^z_{\mu}\delta^z_{\nu} \propto \text{diag}(1, 0, 0, 0)$$

Lowest order

Tree-level amplitude

$$\mathcal{A} \propto G_N \int d^4x \, \frac{T_{\mu\nu}T'^{\mu\nu} - \frac{1}{3}TT'}{\Box - m^2(\Box)} = 0$$

#### Finite size DWs



$$M = \pi L^2 \sigma$$

$$r_* \sim \left(\delta L^2 r_c\right)^{1/3}$$

$$\delta \equiv \frac{\sigma}{M_*^3}$$

Non-linearities do not appear if  $r_* \ll L$ 



 $L \gg \delta r_c$ 

DW probes large distances in longitudinal directions

#### CONCLUSIONS

- In DGP, sub-critical DWs do not gravitate
   => short distance probe of gravity
- expect the same in other IR-modified gravity theories
   (non-linearities do not contribute for DWs)
- Screening mechanism in DGP -> Extrinsic curvature
- Almost the same as for Schwarzchild-like case (Gabadadze&lglesias)
- Super-massive codimension-2 branes
- Screening of the 5D-tension

#### **Domain Walls**

branch	DW tension	$\sigma_{ m eff}^{(4D)}/\sigma$	$\sigma_{ m eff}^{(5D)}/\sigma$
СВ	$\sigma < 2\pi M_*^3$	0	1
	$\sigma > 2\pi M_*^3$	$1 - \frac{2\pi M_*^3}{\sigma}$	$\frac{d}{2r_c}$
SAB	$\sigma \ll 2\pi M_*^3$	$2\left(1-\left(\sigma/M_{*}^{3}\right)^{2}/48+\ldots\right)$	$-\left(1-\left(\sigma/M_{*}^{3}\right)^{2}/24+\ldots\right)$
	$\sigma \gg 2\pi M_*^3$	$1 + \frac{2\pi M_*^3}{\sigma} + \dots$	$-\frac{d}{2r_c}$

# Schwarzchild-DGP

branch	$M^{(4D)}/M$	$M^{(5D)}/M$
СВ	0	$0.56 \; \frac{r_*}{r_c}$
SAB	0	$-0.45 \frac{r_*}{r_c}$

#### **FUTURE**

- notion of  $r_*$  for dilute walls  $(d \ge r_c)$
- less symmetric configurations
- general solutions (curved bulk)
- field theoretical analog??