Ultraviolet properties of maximal supergravity theories

Pierre Vanhove (SPhT – Saclay)

based on some work done in collaboration with M.B. Green and J.Russo

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L-loop amplitudes

A L-loop four gravitons amplitude in D dimensions has the UV behaviour

$$A_L^{(4)}] = \Lambda^{(D-2)L+2}$$

N = 8 higher-loop amplitudes factorize more momenta

$$[A_L^{(4)}] = \Lambda^{(D-2)L-6-2\beta_L} [D^{2\beta_L} R^4]$$

At one-loop one has $\beta_1=0$ and $\beta_2=2$ at two-loops Bern et al., D'Hoker/Phong, Berkovits

In this talk we will give arguments for the validity of

 $\beta_L = L$ for $L \ge 2$

based on the works done with M.B. Green and J. Russo hep-th/0610299, hep-th/0611273, hep-th/0612nnn, hep-th/07mmnnn

Systematics of four gravitons multiloop amplitudes

2 Multiloops amplitudes in eleven dimensions

3 Gravitational F-terms

4 Reduction to Four dimensions

Higher-loop amplitudes factorize some higher power of momenta

$$[A_L^{(4)}] = \Lambda^{(D-2)L-6-2\beta_L} [D^{2\beta_L} R^4]$$

Starting from D = 11 a generic *L*-loop contribution to the operator $D^{2\beta_L} R^4$ after compactification on a circle of radius R_{11} is

$$S_{L}^{(\beta_{L}+\mathbf{v}+3)} = \sum_{w=0}^{w_{L}} \ell_{P}^{9(L-1)} \frac{\Lambda^{9L-6-2\beta_{L}-w}}{R_{11}^{w-2}} (R_{11}^{2}D^{2})^{\mathbf{v}} D^{2\beta_{L}} R^{4}$$
$$S_{L}^{(k+3)} = \sum_{q=\beta_{L}-k}^{q_{L}} \ell_{P}^{9(L-1)} \Lambda^{9L-6-2k-2q} R_{11}^{2(1-q)} D^{2k} R^{4}.$$

The relation between the Type II string and M-theory parameters

$$R_{11}^{3} = (g_{s}^{A})^{2} \ell_{P}^{3}$$

$$ds_{M-th}^{2} = \frac{\ell_{P}^{3}}{\ell_{s}^{2}} R_{11}^{-1} ds_{IIA}^{2} + R_{11}^{2} (dx^{11} - C_{m} dx^{m})^{2}$$

We will assume that these relations always hold and are not corrected by higher order corrections

$$D^2
ightarrow R_{11} \, rac{\ell_P^3}{\ell_s^2} \, D^2$$

UV behaviour of multiloop amplitudes

$$S_{m,h}^{(k+3)A} = \ell_s^{2k-2} \Lambda^{\delta} \int d^{10}x \sqrt{-g_A} c_h e^{2(h-1)\phi_A} D^{2k} R^4$$
$$\delta = 9L - 6 - 2\beta_L - w; \qquad h = 1 + k - \frac{w + 2\beta_L}{3}$$

Since w parametrizes the UV (sub-)divergences of the L-loop diagram

$$w = 3 + 3k - 3h - 2\beta_L \ge 0 \iff h \le k + 1 - \frac{2\beta_L}{3}$$

Since at one-loop $\beta_1 = 0$ and the only diverging term is R^4 Since at higher-loop we have $\beta_L = 2$ for $L \ge 2$ then

$h \leq k$

Non renormalisation conditions of $D^{2k}R^4$ (or R^{4+k}) operators

- There are no contributions with string loop genus h > k.
- The contributions with h = k are determined exactly by finite contributions from the derivative expansion of the one-loop (L = 1) diagram in eleven dimensions.
- Contributions with h < k are permitted and may arise from any number of supergravity loops greater than one (L > 1).

What are the explicit computations in favor of this ?

Relating M-theory to Superstring



• $\mathcal{N}=8$ susy implies that the amplitude factorizes by \mathcal{R}^4 Bern et al.



• We have $\beta_1 = 0$ and degree of divergence $\delta_1 = 3$

• Compactification on T^2 gives type IIA or type IIB string

$$S^{IIA} = \int d^9 x \sqrt{g} \mathcal{R}^4 r_a \left[\frac{2\zeta(3)}{g_s^2} + \frac{4\zeta(2)}{r_a^2} + (\ell_P \Lambda)^3 \right] \\ + \int d^9 x \sqrt{g} r_a \sum_{n \ge 2} \frac{\Gamma(n-1/2)\zeta(2n-1)}{n!} r_a^{2(n-1)} (\alpha' \mathcal{D}^2)^n \mathcal{R}^4 \\ + \int d^9 x \sqrt{g} r_a \frac{\Gamma(n-1)\zeta(2n-2)}{n!} e^{2(n-1)\phi} (\alpha' \mathcal{D}^2)^n \mathcal{R}^4 \\ + \text{ non-anal.} + \text{non-pert.}$$

Relating M-theory to Superstring

T-duality invariance at one-loop 4 graviton IIA amplitude = IIB amplitude

$$r_a \leftrightarrow rac{1}{r_a}; \qquad rac{r_a}{(g_s^a)^2} = rac{r_b}{(g_s^b)^2}$$

$$S = \int d^9 x \sqrt{g} \mathcal{R}^4 \left[r \frac{2\zeta(3)}{g_s^2} + \frac{4\zeta(2)}{r} + r(\ell_P \Lambda)^3 \right]$$

Fixes the value of the cut-off

$$4\zeta(2) = (\ell_P \Lambda)^3$$

The higher-derivative one string loop terms are not T-duality invariant!!

$$S^{IIA} = \int d^9 x \sqrt{g} \sum_{n \ge 2} \frac{\Gamma(n-1/2)\zeta(2n-1)}{n!} r_a^{2n-1} (\alpha' \mathcal{D}^2)^n \mathcal{R}^4$$

T-duality will be recovered once the higher-loop amplitudes contributions are included

Two loops in 11D

• $\mathcal{N}=8$ susy implies that the amplitude factorizes $\mathcal{D}^4\,\mathcal{R}^4$ Bern et al.



• This has $\beta_2 = 2$ and the superficial UV behaviour is $\delta_2 = 8$

Pierre Vanhove (SPhT - Saclay)

Is N = 8 supergravity finite?

Two-loop in 11d

• The planar and non-planar φ^3 double box amplitude sum into a modular invariant amplitude



$$\mathcal{A}^{P+NP} = \mathcal{D}^{4} \mathcal{R}^{4} \int_{0}^{\Lambda^{2}} dV V^{3} \int_{\Gamma_{0}(2)} \frac{d^{2}\tau}{\tau_{2}^{2}} f(V,\tau;s,t,u)$$
$$V = \ell_{P}^{2} (\sigma\rho + \sigma\lambda + \rho\lambda)^{-1/2}$$

• Gives a uniform way of putting the cut-off

Two-loop in 11d



• The planar and non-planar φ^3 double box amplitude sum into a modular invariant amplitude

$$\begin{aligned} \mathcal{A}^{P+NP} &= \\ \mathcal{D}^4 \mathcal{R}^4 \int_0^{\Lambda^2} dV V^3 \int_{\Gamma_0(2)} \frac{d^2 \tau}{\tau_2^2} f(V,\tau;s,t,u) \\ V &= \ell_P^2 (\sigma \rho + \sigma \lambda + \rho \lambda)^{-1/2} \end{aligned}$$

• Gives a uniform way of putting the cut-off

The 2-loop amplitude on \mathcal{T}^2 of volume $\mathcal V$ and complex structure Ω

$$\mathcal{A}^{P+NP}(\mathcal{V},\Omega) = \int_{0}^{\Lambda^{2}} dV V^{3} \int_{\Gamma_{0}(2)} \frac{d^{2}\tau}{\tau_{2}^{2}} \Gamma_{(2,2)}(\mathcal{V},\Omega;V,\tau) f(V,\tau;s,t,u)$$

$$f(V,\tau;s,t,u) = D^{4}R^{4} + \frac{A_{1}(\tau,\bar{\tau})}{V} D^{6}R^{4} + \frac{A_{2}(\tau,\bar{\tau})}{V^{2}} D^{8}R^{4} + \cdots$$

The coupling $A_{(n)}(\tau, \bar{\tau}) = \sum_{i} a^{i}_{(n)}(\tau, \bar{\tau})$ satisfies

$$\Delta_{\Omega} a_{(n)}^{i} = \lambda_{i} a_{(n)}^{i} + P_{N}(\tau_{2}) \delta(\tau_{1})$$

Source term gets contributions from the cusps of $\Gamma_0(2)$

Implies that the 10d coupling satisfies

$$I_{D^{4}R^{4}} = \cdots + \frac{(\ell_{P}\Lambda)^{3}}{\mathcal{V}^{5}} E_{5/2} + \cdots$$
$$I_{D^{6}R^{4}} = \cdots + \frac{1}{\mathcal{V}^{3}} E_{(3/2,3/2)} + \cdots$$

where

$$\Delta_{\Omega} E_{5/2} = \frac{15}{4} E_{5/2}$$
$$\Delta_{\Omega} E_{(3/2,3/2)} = 12 E_{(3/2,3/2)} - 6 E_{3/2}^2$$

• Type IIB theory is obtained by taking M-theory on a 2-torus of vanishing volume \mathcal{V} with fixed complex structure $\Omega \equiv \tau$

$$S = \int d^9 x \sqrt{-G^{(9)}} \mathcal{V} \sum_n h^{(n)}(\mathcal{V}, \Omega) D^{2n} R^4$$
$$h^{(n)}(\mathcal{V}, \Omega) = \cdots + \frac{1}{\mathcal{V}^{\frac{3+n}{2}}} f^{(n)}(\Omega) + \cdots$$

• Only this contribution decompactifies to the 10d type limit

$$S = \int d^{10}x \sqrt{-g^{(10)}} e^{\frac{n-1}{2}\phi} f^{(n)}(\tau) D^{2n} R^4$$

Two loops in 11D

- The $D^4 R^4$ coupling is given by the $E_{5/2}$ series and gets tree-level, no one-loop contributions and a two-loop superstring contributions The perturbative results were checked by Green/Vanhove;D'Hoker/Phong.
- The $D^6 R^4$ is given by the generalized series $E_{(3/2,3/2)}$ and gets tree-level, one-, two- and three-loop superstring contributions. The one-loop result was confirmed by Green/Vanhove. The 3-loop coefficient matches the one predicts one the type IIA side by the L = 1 amplitude.

The F-term computation of Berkovits shows that this contribution is not renormalised after 3-loop.

These ten dimensional couplings should not get corrections from higher-loop amplitudes in 11D

Three loops in 11D



• The ladder diagram has $\beta_3 = 4$ and Mondrian diagram have $\beta_3 = 2$ Bern et al.

Three loops in 11D (one-loop subdivergences)



Gravitational F-terms

Using the non-minimal pure spinor formalism Berkovits showed that the $D^{2h}R^4$ up to h = 5 are **F-terms**

At one-loop

$$A_1^{(4)}\sim\int d^{16} heta_L d^{16} heta_R\, heta_L^{11} heta_R^{11}\,AW^3\sim R^4$$

At higher-loop up to genus h = 5

$$\begin{array}{ll} A_h^{(4)} & \sim & \int |d\mu_h|^2 \int d^{16}\theta_L d^{16}\theta_R \, \theta_L^{12-2h} \theta_R^{12-2h} \, W^4 \, |(dx)^{g-2}|^2 \, e^{ik \cdot X} \\ \\ & \sim & \int |d\mu_h|^2 \, |(dx)^{h-2}|^2 \, D^{2h} R^4 \, e^{ik \cdot X} \end{array}$$

where

$$W_{\alpha\beta} = F_{\alpha\beta} + \dots + \theta_L^{\gamma} \theta_R^{\delta} R_{\alpha\gamma\beta\delta} + \dots$$
$$A_{\alpha\beta} = \theta_L^{\gamma} \theta_R^{\delta} g_{\alpha\gamma\beta\delta} + \dots$$

In ten-dimensions $\beta_L = L$ for $2 \le L \le 5$ and $\beta_L \ge 6$ for $L \ge 6$

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When

$$\beta_L = L$$

the amplitude have a reduced UV behaviour since for

$$A_L^{(4)}\sim D^{2L}R^4$$
 I(s, t, u; $lpha'/R^2$)

the residual loop integrals $I(s, t, u; \alpha' R)$ behaves as

$$\delta_L = (D-4)L - 6$$

typical of φ^4 or φ^3 coupled to electromagnetism

Ultraviolet behaviour in N = 8 supergravity

• The reduction to D = 4 with $\beta_L = L$ leads to

$$A_L^{(4)} \sim D^{2L} R^4 I(s, t, u; \alpha'/R^2)$$

• where I has to UV power counting behaviour

$$\delta_L = -6 < 0$$
 for $L \ge 2$

- So there is no UV divergences
- But the amplitude can have IR divergences

Ultraviolet behaviour in N = 8 supergravity

• The low-energy limit of string theory is obtained by taking

$$\alpha' \to 0; \qquad \kappa_{(4)}^2 = \frac{g_s^2}{V_6} {\alpha'}^4 = \text{fixed}$$

• in this limit NS-branes become massless

$$E_{NS5} = \frac{\alpha'}{\kappa_{(4)}^2}$$

• They could be needed to cure the IR singularities of the amplitude